Design, testing and preliminary results of the Antarctic infrared binocular telescope (AIRBT)

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Design

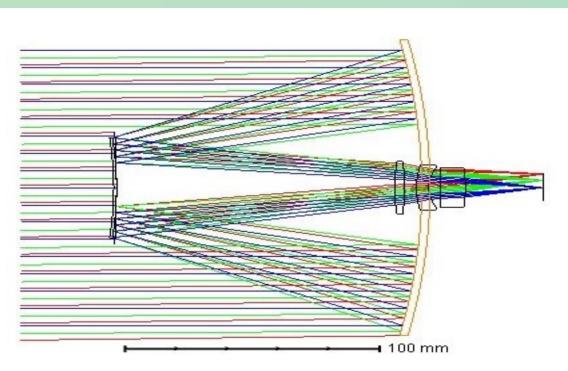


Figure 1: Optical path.

Telescope

Diameter	$15 \mathrm{cm} \times 2$
f/	3
Band	J,H',(W)

Camera

Size	$640 \times 512@15 \mu \mathrm{m}$
Band	$0.9{\sim}1.7\mu{ m m}$
QE	$\geq 70\%$
Dark	$613\mathrm{e^{-}/s/p}$ @- $55^{\circ}\mathrm{C}$
Full well	$100 \mathrm{ke}^-$
Rdnoise	$134e^{-}(Nom.50e^{-})$
Nonlinear	$\leq 0.2\%$

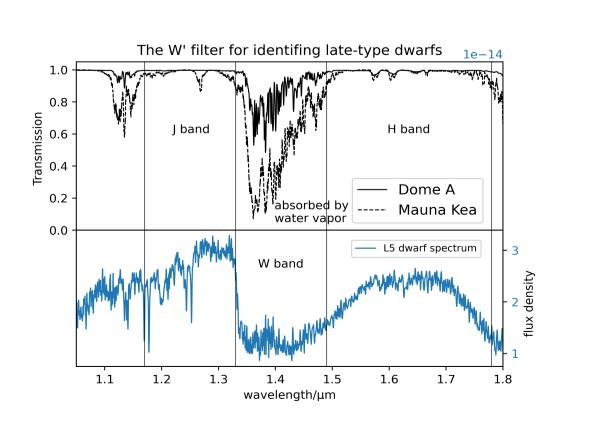


Figure 2: W band filter design (Sims et al. 2012; MaunaKea.com).

Testing

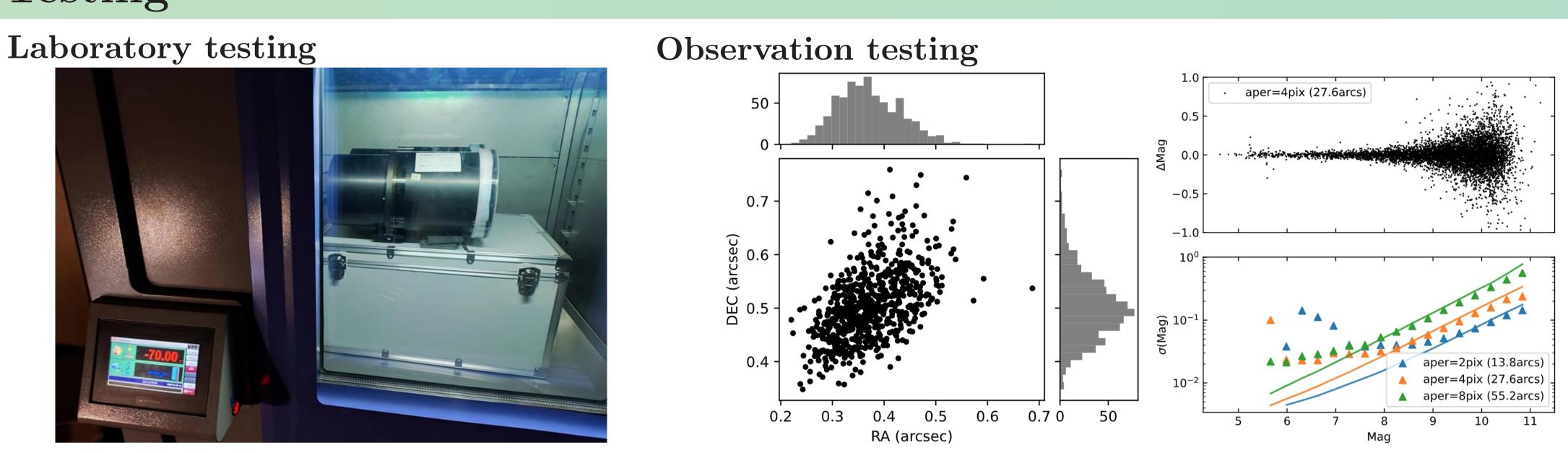


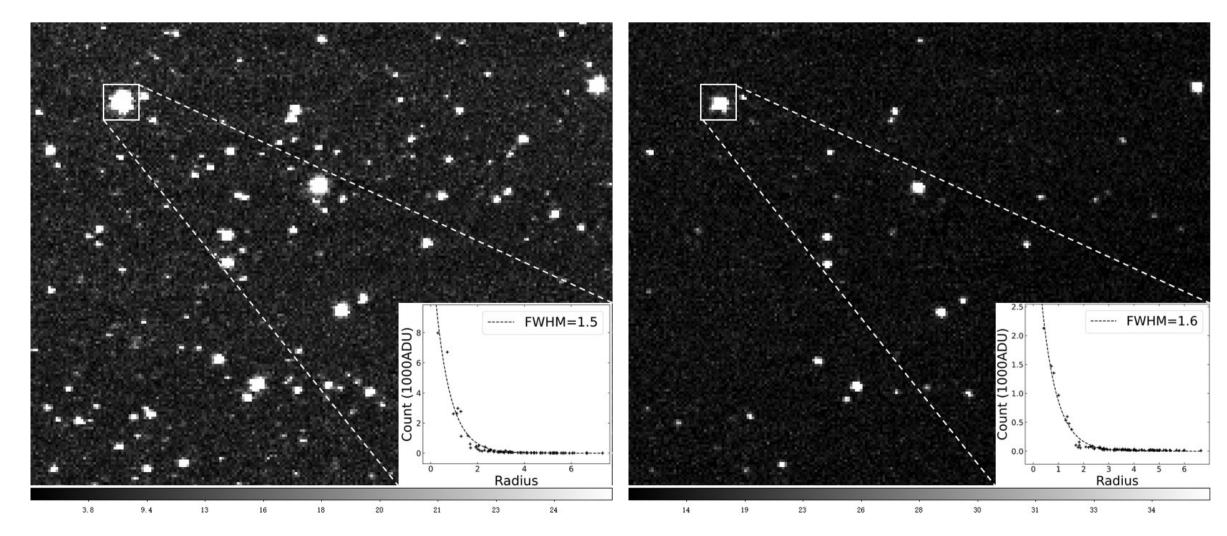
Figure 3: We tested the cold start and hot ru- Figure 4: We observation in Zhuhai City using 3s middle mode. Astrom--nning of the equipment in a cryogenic chamber.etry by Scamp (left), and photometry by Sextractor (right). Photometry accuracy is 21mmag.

Preliminary results

The AIRBT aimed at monitoring the infrared site condition and conducting research in time-domain astronomy, such as transient like GRB and variable stars. The AIRBT project will provide references and accumulate technical experience for the deployment of larger-aperture near-infrared telescopes in Antarctica. More observations and data processing are underway. Welcome cooperation!



Figure 5: Installed in 8m tower in DomeA by China 39th, 40th Antarctic Expedition team.



12.5 in J band (left) and 10 in H band (right).

Figure 6: 1/16 of $1^{\circ} \times 1^{\circ} . 2$ FOV stack $2s \times 30$ frames. Limiting magnitude